Hot Gas & Dark Matter Far Out in the Halos of Massive Galaxies and Groups

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Outline

Missing Baryons & Feedback

- Isolated Elliptical Galaxy: NGC 720
- Fossil Group/Cluster: RXJ 1159+5531

Tilt of the Fundamental Plane

• Mass Slope -- Size Relation for Ellipticals

Missing Baryons: Hot Gas Around Spirals

(Benson, Bower, Frenk, & White 2000, MNRAS, 314, 557)



(Rasmussen et al. 2009)

•No evidence for large amount of extended hot gas

"Missing Baryons"
not yet found in
disk galaxies

(Rasmussen et al. 2009)



X-Ray Constraints on Masses of Ellipticals



(Humphrey, Buote, Gastaldello, Zappacosta, Bullock, Brighenti, & Mathews 2006, ApJ, 646, 899)

NGC 720

- •D=26 Mpc
- •E4

•Very Isolated, No AGN disturbance



(Humphrey, Buote, Canizares, Fabian, & Miller 2011, ApJ, 729, 53)







Mass Determination Method

Input S(r) & M(r)
Solve
$$\oint$$
 H.E.

Output
$$\rho_{gas}(r)$$
 and $T(r)$

We use assumed parameterized models of S & M and then fit the ρ_{gas} & T data

Free Parameters

- Temperature/Pressure -- normalization (1)
- $S \propto \rho^{-2/3}T$ -- broken power-law + const (5)
- M (3)
 - Black Hole -- M_{BH} (fixed)
 - DM Halo -- NFW C Δ , M Δ (2)
 - Stars -- $M_{\star}/L_{J}(I)$

9 Free Parameters Constrained by 26 Data Points









Group and Cluster Data from:

Gastaldello et al. 2007; Sun et al. 2009; Vikhlinin et al. 2006; Dai et al. 2009

Baryon fractions from Giodini et al. 2009

Future Work

- Deep, Offset Suzaku Data for N720 to go out to R₅₀₀
- Follow-up of X-ray snapshot sample of 35 isolated ellipticals -- w/T. Ponman, E. O'Sullivan (Birmingham) & C.Topchyan (UC Irvine)
- Combine X-rays & Stellar Dynamics

Non-Thermal Gas Support in NGC 4649



(Humphrey, Buote, Gebhardt, Brighenti, & Mathews, in prep)

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Baryon Fractions of Galaxy Groups



(Gastaldello, Buote, Humphrey, Zappacosta, Bullock, Brighenti, & Mathews 2007, ApJ, 669, 158)

RXJ 1159+5531

- •Fossil Group/Cluster (~10¹⁴ M_{sun})
- •Redshift 0.08
- •Virial Radius (R100) fits on single Suzaku pointing

(Humphrey, Buote, Brighenti, Flohic, Gastaldello, & Mathews, in prep)





$R_{vir} = 1100 \text{ kpc}$





Tuesday, March 1, 2011



Virgo Cluster



(Simionescu et al. 2011, arXiv:1102.2429)



Future Work

- Additional targets with virial radius observed in single Suzaku obs
- Systematic study of I5 X-ray--selected Groups

Power-Law Mass Profiles



(Humphrey & Buote 2010, MNRAS, 403, 2143)

Slope-R_e Relation



(Humphrey & Buote 2010, MNRAS, 403, 2143)

Similar α -R_e relation confirmed by SLACS (Auger et al. 2010, ApJ, 724, 511)

Tilt of Fundamental Plane



Are Nearly Power-Law Mass Profiles a Fundamental Feature of Galaxy Formation?